1 MLE and MAP for Regression (Part II)

So far we have used MLE and MAP to justify the optimization formulation of OLS and ridge regression, respectively. The MLE formulation assumes that the observation $Y_i$ is a noisy version of the true underlying output:

$$Y_i = f(x_i) + Z_i$$

where the noise for each datapoint is i.i.d. The MAP formulation assumes that the model parameter $W_j$ is according to an i.i.d. Gaussian prior

$$W_j \overset{iid}{\sim} \mathcal{N}(\mu_j, \sigma^2_h)$$

So far, we have restricted ourselves to the case when the noise/parameters are i.i.d:

$$Z \sim \mathcal{N}(0, \sigma^2 I), \quad W \sim \mathcal{N}(\mu_W, \sigma^2_h I)$$

However, what about the case when $N_i$’s/$W_j$’s are non-identical or dependent on one another? We would like to explore the case when the noise and parameters are jointly Gaussian with arbitrary covariance matrices:

$$Z \sim \mathcal{N}(0, \Sigma_Z), \quad W \sim \mathcal{N}(\mu_W, \Sigma_W)$$

It turns out that via a change of coordinates, we can reduce these non i.i.d. problems back to the i.i.d. case and solve them using the original techniques we used to solve OLS and Ridge Regression!

1.1 Weighted Least Squares

The basic idea of **weighted least squares** is the following: we place more emphasis on the loss contributed from certain data points over others - that is, we care more about fitting some data points over others.

1.1.1 Optimization View

From an optimization perspective, the problem can be expressed as

$$w^*_{WLS} = \arg \min_{w \in \mathbb{R}^d} \left( \sum_{i=1}^{n} \omega_i (y_i - x_i^T w)^2 \right)$$
This objective is the same as OLS, except that each term in the sum is weighted by a positive coefficient $\omega_i$. As always, we can vectorize this problem:

$$w_{WLS}^* = \arg \min_{w \in \mathbb{R}^d} (y - Xw)^\top \Omega (y - Xw)$$

Where the $i$'th row $X$ is $x_i^\top$, and $\Omega \in \mathbb{R}^{n \times n}$ is a diagonal matrix with $\Omega_{i,i} = \omega_i$.

We rewrite the WLS objective to an OLS objective:

$$w_{WLS}^* = \arg \min_{w \in \mathbb{R}^d} (y - Xw)^\top \Omega (y - Xw) = \arg \min_{w \in \mathbb{R}^d} (y - Xw)^\top \Omega^{1/2} \Omega^{1/2} (y - Xw)$$

$$= \arg \min_{w \in \mathbb{R}^d} (\Omega^{1/2} y - \Omega^{1/2} Xw)^\top (\Omega^{1/2} y - \Omega^{1/2} Xw)$$

$$= \arg \min_{w \in \mathbb{R}^d} \|\Omega^{1/2} y - \Omega^{1/2} Xw\|^2$$

This formulation is identical to OLS except that we have scaled the data matrix and the observation vector by $\Omega^{1/2}$, and we conclude that

$$w_{WLS}^* = \left( (\Omega^{1/2} X)^\top (\Omega^{1/2} X) \right)^{-1} \left( \Omega^{1/2} X \right)^\top \Omega^{1/2} y = (X^\top \Omega X)^{-1} X^\top \Omega y$$

1.1.2 Probabilistic View

As in MLE, we assume that our observations $y$ are noisy, but now suppose that some of the $y_i$'s are more noisy than others. How can we take this into account in our learning algorithm so we can get a better estimate of the weights? Our probabilistic model looks like

$$Y_i = x_i^\top w + Z_i$$

where the $Z_i$'s are still independent Gaussians random variables, but not necessarily identical: $Z_i \sim \mathcal{N}(0, \sigma_i^2)$. Jointly, we have that $Z \sim \mathcal{N}(\mu_Z, \Sigma_Z)$, where

$$\Sigma_Z = \begin{bmatrix} \sigma_1^2 & 0 & \cdots & 0 \\ 0 & \sigma_2^2 & \cdots & 0 \\ \vdots & \vdots & \ddots & \vdots \\ 0 & \cdots & \cdots & \sigma_n^2 \end{bmatrix}$$

We can morph the problem into an MLE one by scaling the data to make sure all the $Z_i$'s are identically distributed, by dividing by $\sigma_i$:

$$\frac{Y_i}{\sigma_i} = \frac{x_i^\top}{\sigma_i} w + \frac{Z_i}{\sigma_i}$$

Note that the scaled noise entries are now i.i.d:

$$\frac{Z_i}{\sigma_i} \sim \mathcal{N}(0, 1)$$
Jointly, we can express this change of coordinates as

\[ \Sigma^{-\frac{1}{2}}y \sim N(\Sigma^{-\frac{1}{2}}Xw, \Sigma^{-\frac{1}{2}}\Sigma Z_\Sigma^{-\frac{1}{2}}) = N(\Sigma^{-\frac{1}{2}}Xw, I) \]

This change of variable is sometimes called the **reparameterization trick**. Now that the noise is i.i.d. using the change of coordinates, we rewrite our original problem as a scaled MLE problem:

\[
w_{\text{WLS}}^* = \arg \min_{w \in \mathbb{R}^d} \left( \sum_{i=1}^{n} \frac{(y_i - \frac{x_i^\top}{\sigma_i} w)^2}{2} \right) + n \log \sqrt{2\pi}
\]

\[
= \arg \min_{w \in \mathbb{R}^d} \sum_{i=1}^{n} \frac{1}{\sigma_i^2} (y_i - x_i^\top w)^2
\]

The MLE estimate of this scaled problem is equivalent to the WLS estimate of the original problem:

\[
w_{\text{WLS}}^* = (X^\top \Sigma^{-\frac{1}{2}} Z_\Sigma^{-\frac{1}{2}} X)^{-1} X^\top \Sigma^{-\frac{1}{2}} Z_\Sigma^{-\frac{1}{2}} y = (X^\top \Sigma^{-1} X)^{-1} X^\top \Sigma^{-1} y
\]

As long as no \( \sigma \) is 0, \( \Sigma Z \) is invertible. Note that \( \omega_i \) from the optimization perspective is directly related to \( \sigma_i^2 \) from the probabilistic perspective: \( \omega_i = \frac{1}{\sigma_i^2} \). As the variance \( \sigma_i^2 \) of the noise corresponding to data point \( i \) decreases, the weight \( \omega_i \) increases: we are more concerned about fitting data point \( i \) because it is likely to match the true underlying denoised point. Inversely, as the variance \( \sigma_i^2 \) increases, the weight \( \omega_i \) decreases: we are less concerned about fitting data point \( i \) because it is noisy and should not be trusted.

### 1.2 Least Squares with Dependent Noise

Now let's consider the case when the noise random variables are dependent on one another. We have

\[ Y = Xw + Z \]

where \( Z \) is now a jointly Gaussian random vector. That is,

\[ Z \sim \mathcal{N}(0, \Sigma_Z), \quad Y \sim \mathcal{N}(Xw, \Sigma_Z) \]

Our goal is to maximize the probability of our data over the set of possible \( w \)’s:

\[
w^* = \arg \max_{w \in \mathbb{R}^d} \frac{1}{\sqrt{\det(\Sigma_Z)}} \frac{1}{(\sqrt{2\pi})^n} e^{-\frac{1}{2}(y-Xw)^\top \Sigma_Z^{-1}(y-Xw)}
\]

\[
= \arg \min_{w \in \mathbb{R}^d} (y - Xw)^\top \Sigma_Z^{-1}(y - Xw)
\]

Since \( \Sigma_Z \) is symmetric, we can decompose it into its eigen structure using the spectral theorem:

\[
\Sigma_Z = Q \begin{bmatrix} \sigma_1^2 & 0 & \cdots & 0 \\ 0 & \sigma_2^2 & \cdots & 0 \\ \vdots & \vdots & \ddots & \vdots \\ 0 & \cdots & \cdots & \sigma_n^2 \end{bmatrix} Q^\top
\]
where $Q$ is orthonormal. As before with weighted least squares, our goal is to find an appropriate linear transformation so that we can reduce the problem into the i.i.d. case.

Consider

$$
\Sigma_{Z}^{-\frac{1}{2}} = Q \begin{bmatrix}
\frac{1}{\sigma_1} & 0 & \cdots & 0 \\
0 & \frac{1}{\sigma_2} & \cdots & 0 \\
\vdots & \vdots & \ddots & \vdots \\
0 & \cdots & \cdots & \frac{1}{\sigma_n}
\end{bmatrix} Q^\top
$$

We can scale the data to morph the problem into an MLE problem with i.i.d. noise variables, by premultiplying the data matrix $X$ and the observation vector $y$ by $\Sigma_{Z}^{-\frac{1}{2}}$. Jointly, we can express this change of coordinates as

$$
\Sigma_{Z}^{-\frac{1}{2}} y \sim \mathcal{N}(\Sigma_{Z}^{-\frac{1}{2}} X w, \Sigma_{Z}^{-\frac{1}{2}} \Sigma_{Z} \Sigma_{Z}^\top \Sigma_{Z}^{-\frac{1}{2}}) = \mathcal{N}(\Sigma_{Z}^{-\frac{1}{2}} X w, I)
$$

In a very similar fashion to the independent noise problem, the MLE of the scaled dependent noise problem is

$$
w^* = (X^\top \Sigma_{Z}^{-\frac{1}{2}} X)^{-1} X^\top \Sigma_{Z}^{-\frac{1}{2}} y
$$

### 1.3 Ridge Regression with Dependent Parameters

In the ordinary least squares (OLS) statistical model, we assume that the output $Y$ is a linear function of the input, plus some Gaussian noise. We take this one step further in MAP estimation, where we assume that the weights are a random variable. The new statistical model is

$$
Y = X W + Z
$$

where $Y$ and $Z$ are $n$-dimensional random vectors, $W$ is a $d$-dimensional random vector, and $X$ is a fixed $n \times d$ matrix. Note that random vectors are not notationally distinguished from matrices here, so keep in mind what each symbol represents.

We have seen that ridge regression can be derived by assuming a prior distribution on $W$ in which $W_i$ are i.i.d. (univariate) Gaussian, or equivalently,

$$
W \sim \mathcal{N}(0, I)
$$

But more generally, we can allow $W$ to be any multivariate Gaussian:

$$
W \sim \mathcal{N}(\mu_W, \Sigma_W)
$$

Recall that we can rewrite a multivariate Gaussian variable as an affine transformation of a standard Gaussian variable:

$$
W = \Sigma_{W}^{1/2} V + \mu_W \quad \text{V} \sim \mathcal{N}(0, I)
$$

Plugging this parameterization into our previous statistical model gives

$$
Y = X (\Sigma_{W}^{1/2} V + \mu_W) + Z
$$
But this can be re-written
\[ Y - X\mu_W = X\Sigma_W^{1/2}V + Z \]
which we see has the form of the statistical problem that underlies OLS, and therefore
\[ \hat{v} = (\Sigma_W^{1/2}X^TX\Sigma_W^{1/2} + I)^{-1}\Sigma_W^{1/2}X^T(y - X\mu_W) \]

However \( V \) is not what we care about – we need to convert back to the actual weights \( W \) in order to make predictions. Since \( W \) is completely determined by \( V \) (assuming fixed mean and covariance),
\[ \hat{w} = \Sigma_W^{1/2}\hat{v} + \mu_W \]
\[ = \mu_W + \Sigma_W^{1/2}(\Sigma_W^{1/2}X^TX\Sigma_W^{1/2} + I)^{-1}\Sigma_W^{1/2}X^T(y - X\mu_W) \]
\[ = \mu_W + (X^TX + \Sigma_W^{-1/2}\Sigma_W^{-1/2})^{-1}X^T(y - X\mu_W) \]

Note that there are two terms: the prior mean \( \mu_W \), plus another term that depends on both the data and the prior. The precision matrix of \( W \)'s prior (\( \Sigma_W^{-1} \)) controls how the data fit error affects our estimate.

To gain intuition, let us consider the simplified case where
\[ \Sigma_W = \begin{bmatrix} \sigma_1^2 & 0 & \cdots & 0 \\ 0 & \sigma_2^2 & \cdots & 0 \\ \vdots & \vdots & \ddots & \vdots \\ 0 & 0 & \cdots & \sigma_n^2 \end{bmatrix} \]
When the prior variance \( \sigma_j^2 \) for dimension \( j \) is large, the prior is telling us that \( W_j \) may take on a wide range of values. Thus we do not want to penalize that dimension as much, preferring to let the data fit sort it out. And indeed the corresponding entry in \( \Sigma_W^{-1} \) will be small, as desired.
Conversely if \( \sigma_j^2 \) is small, there is little variance in the value of \( W_j \), so \( W_j \approx \mu_j \). As such we penalize the magnitude of the data-fit contribution to \( \hat{W}_j \) more heavily.

### 1.3.1 Alternative derivation

MAP with colored noise can be expressed as:
\[ U, V \overset{iid}{\sim} \mathcal{N}(0, I) \]
\[ \begin{bmatrix} Y \\ W \end{bmatrix} = \begin{bmatrix} R_Z & XR_W \\ 0 & R_W \end{bmatrix} \begin{bmatrix} U \\ V \end{bmatrix} \]
where \( R_Z \) and \( R_W \) are relationships with \( W \) and \( Z \), respectively. Note that the \( R_W \) appears twice because our model assumes \( Y = XW + \text{noise} \), so if \( W = R_WV \), then we must have \( Y = XR_WV + \text{noise} \).
We want to find the posterior \( W \mid Y = y \). The formulation above makes it relatively easy to find the posterior of \( Y \) conditioned on \( W \) (see below), but not vice-versa. So let’s pretend instead that

\[
\begin{bmatrix}
U' \\
V'
\end{bmatrix} \overset{\text{iid}}{\sim} \mathcal{N}(0, I)
\]

\[
\begin{bmatrix}
W \\
Y
\end{bmatrix} = \begin{bmatrix}
A & B \\
0 & D
\end{bmatrix} \begin{bmatrix}
U' \\
V'
\end{bmatrix}
\]

Now \( W \mid Y = y \) is straightforward. Since \( V' = D^{-1}Y \), the conditional mean and variance of \( W \mid Y = y \) can be computed as follows:

\[
E(W \mid Y = y) = E[AU' + BV' \mid Y = y]
= E[AU' \mid Y = y] + E[BD^{-1}Y \mid Y = y]
= A E[U'] + E[BD^{-1}Y \mid Y = y]
= BD^{-1}y
\]

\[
\text{Var}(W \mid Y = y) = E[(W - E[W])(W - E[W])^\top \mid Y = y]
= E[(AU' + BD^{-1}Y - BD^{-1}Y)(AU' + BD^{-1}Y - BD^{-1}Y)^\top \mid Y = y]
= E[(AU')(AU')^\top \mid Y = y]
= E[AU'(U')^\top A^\top]
= A E[U'(U')^\top] A^\top
= AA^\top
\]

In both cases above where we drop the conditioning on \( Y \), we are using the fact \( U' \) is independent of \( V' \) (and thus independent of \( Y = DV' \)). Therefore

\[
W \mid Y = y \sim \mathcal{N}(BD^{-1}Y, AA^\top)
\]

Recall that a Gaussian distribution is completely specified by its mean and covariance matrix. We see that the covariance matrix of the joint distribution is

\[
E \begin{bmatrix}
W \\
W^\top
\end{bmatrix} = \begin{bmatrix}
A & B \\
0 & D
\end{bmatrix} \begin{bmatrix}
A^\top & 0 \\
B^\top & D^\top
\end{bmatrix}
= \begin{bmatrix}
AA^\top + BB^\top & BD^\top \\
DB^\top & DD^\top
\end{bmatrix}
= \begin{bmatrix}
\Sigma_W & \Sigma_{W,Y} \\
\Sigma_{Y,W} & \Sigma_Y
\end{bmatrix}
\]

Matching the corresponding terms, we can express the conditional mean and variance of \( W \mid Y = y \) in terms of these (cross-)covariance matrices:

\[
BD^{-1}Y = BD^{-1}D^{-1}Y = (BD^\top)(DD^\top)^{-1}Y = \Sigma_{W,Y} \Sigma_Y^{-1}Y
\]
\[
AA^\top = AA^\top + BB^\top - BB^\top \\
= AA^\top + BB^\top - B D^\top D^{-1} D B^\top \\
= AA^\top + BB^\top - (BD^\top)(DD^\top)^{-1}DB^\top \\
= \Sigma_W - \Sigma_{W,Y} \Sigma_Y^{-1} \Sigma_{Y,W}
\]

We can then apply the same reasoning to the original setup:

\[
E \begin{bmatrix} Y \\ W \end{bmatrix} \begin{bmatrix} Y^\top \\ W^\top \end{bmatrix} = \begin{bmatrix} R_Z R_Z^\top + X R_W R_W^\top X^\top & XR_W R_W^\top \\ R_W R_W^\top X^\top & R_W R_W^\top \end{bmatrix}
\]

\[
= \begin{bmatrix} \Sigma_Y & \Sigma_{Y,W} \\ \Sigma_{W,Y} & \Sigma_W \end{bmatrix}
\]

Therefore after defining \( \Sigma_Z = R_Z R_Z^\top \), we can read off

\[
\Sigma_W = R_W R_W^\top \\
\Sigma_Y = \Sigma_Z + X \Sigma_W X^\top \\
\Sigma_{Y,W} = X \Sigma_W \\
\Sigma_{W,Y} = \Sigma_W X^\top
\]

Plugging this into our estimator yields

\[
\hat{w} = E[W \mid Y = y] \\
= \Sigma_{W,Y} \Sigma_Y^{-1} y \\
= \Sigma_W X^\top (\Sigma_Z + X \Sigma_W X^\top)^{-1} y
\]

One may be concerned because this expression does not take the form we expect – the inverted matrix is hitting \( y \) directly, unlike in other solutions we’ve seen. But using the Woodbury matrix identity\(^1\), it turns out that we can rewrite this expression as

\[
\hat{w} = (X^\top \Sigma_Z^{-1} X + \Sigma_W^{-1})^{-1} X^\top \Sigma_Z^{-1} y
\]

which looks more familiar.

### 1.4 Summary of Linear Gaussian Statistical Models

We have seen a number of related linear models, with varying assumptions about the randomness in the observations and the weights. We summarize these below:

<table>
<thead>
<tr>
<th>Z ( W )</th>
<th>( \mathcal{N}(0, I) )</th>
<th>( \mathcal{N}(0, \Sigma_Z) )</th>
</tr>
</thead>
<tbody>
<tr>
<td>No prior</td>
<td>( \hat{w}_{\text{OLS}} = (X^\top X)^{-1} X^\top y )</td>
<td>( \hat{w}_{\text{WLS}} = (X^\top \Sigma_Z^{-1} X + \Sigma_W^{-1})^{-1} X^\top \Sigma_Z^{-1} y )</td>
</tr>
<tr>
<td>( \mathcal{N}(0, \lambda^{-1} I) )</td>
<td>( \hat{w}_{\text{RIDGE}} = (X^\top X + \lambda I)^{-1} X^\top y )</td>
<td></td>
</tr>
<tr>
<td>( \mathcal{N}(\mu_W, \Sigma_W) )</td>
<td>( \hat{w} = \mu_W + (X^\top X + \Sigma_W^{-1})^{-1} X^\top (y - X \mu_W) )</td>
<td></td>
</tr>
</tbody>
</table>

\(^1\) \( (A + UCV)^{-1} = A^{-1} - A^{-1} U (C^{-1} + V A^{-1} U)^{-1} V A^{-1} \)